Solar Storm Modeling using OpenACC: From HPC cluster to "in-house"

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Predictive Science Inc. www.predsci.com



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Outline

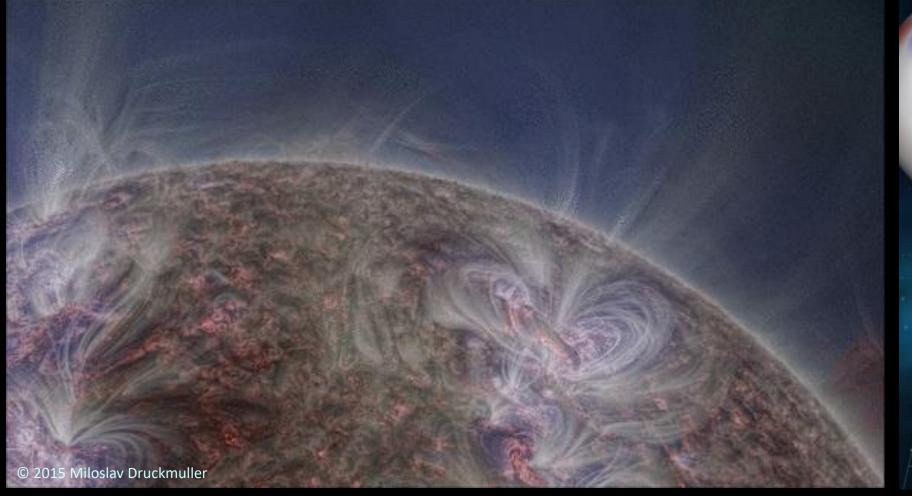
- Solar Storms
- Modeling a Coronal Mass Ejection
- Why add OpenACC?
- Recap of previous OpenACC implementations
- MAS: Magnetohydrodynamic Algorithm outside a Sphere
- Initial OpenACC Implementation of MAS
- "Time-to-solution" results
- Summary and Outlook

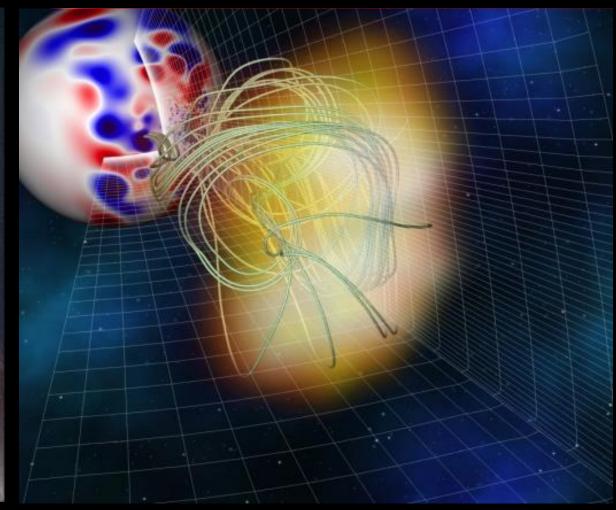


Solar Storms

- Solar storms include coronal mass ejections (CMEs): large explosive events capable of ejecting a billion tons of magnetized million-degree plasma out into space
- CME impacts on Earth can cause interference and damage to electronic infrastructure including GPS satellites and the power grid
- The first step in forecasting CME impacts is the ability to accurately model their initiation and propagation





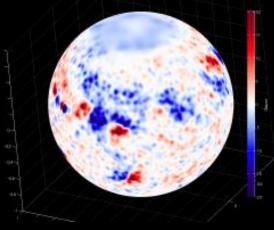


How We Model a Coronal Mass Ejection

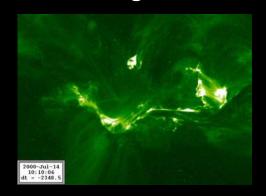
Observations



Satellite Observations



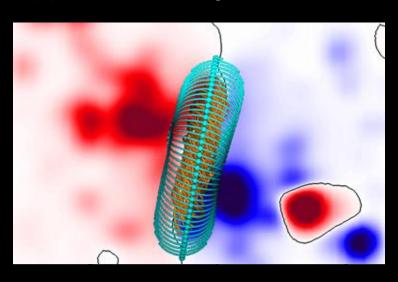
Surface Magnetic Field

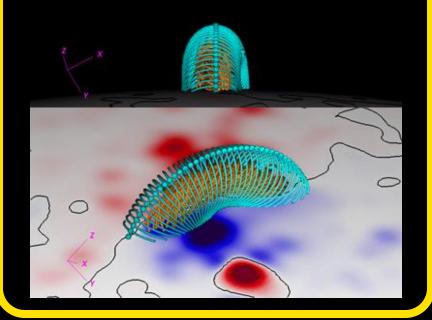


EUV images

CME Initial Condition

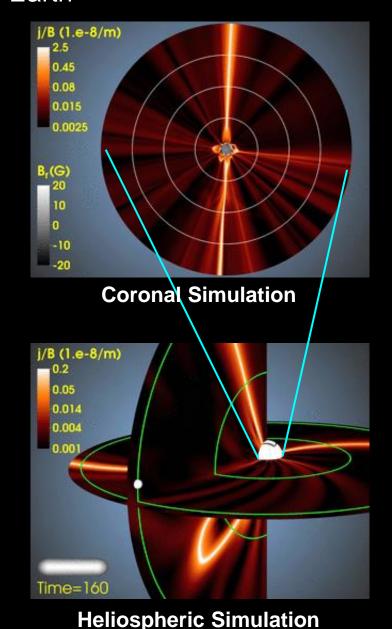
Design and compute stable "Flux Rope" in "Active Region" embedded in global approximate magnetic field



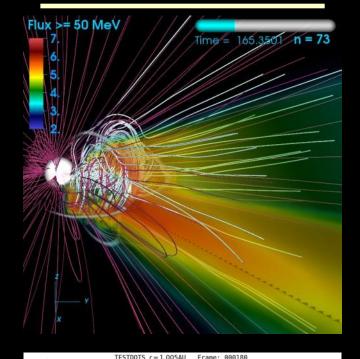


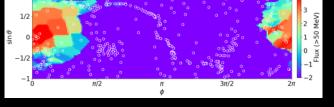
Global TMHD Simulations

Manipulate surface field/flow to erupt CME and propagate to Earth

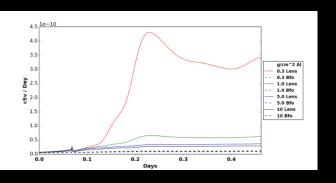


Post Analysis





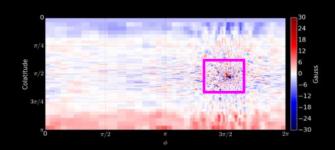
Energetic Particle Fluxes



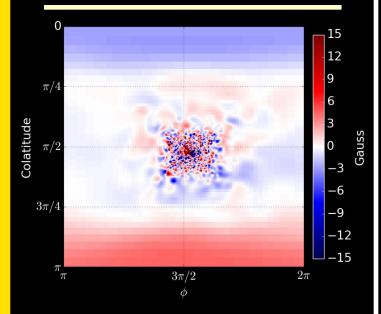
Radiation Dose Levels

Flux Rope Modeling Pipeline (CME Generator)

Isolate CME location, set grid and interpolate



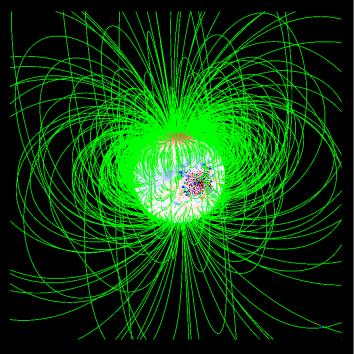
Smooth data to resolve





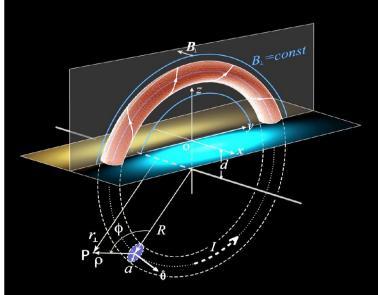
Compute approximate 3D magnetic field

Potential Field

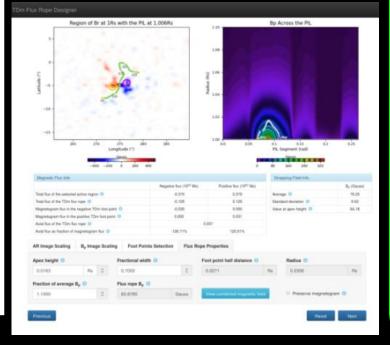




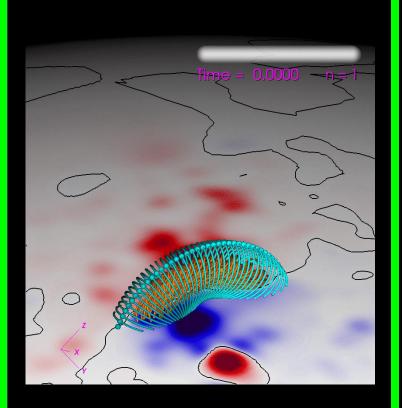
Design and insert analytic flux rope



Titov, V.S., et. al. Ap.J. 790,163 (2014)



Relax to Steady-State with "0-Beta" MHD Simulation





Production Test Run (TEST1)

TEST1: Stable rope (default resolution)

Run information

Physical time duration: **211 sec**

Number of time-steps: 200

 $160 \times 267 \times 602 \sim 26 \text{ million points}$

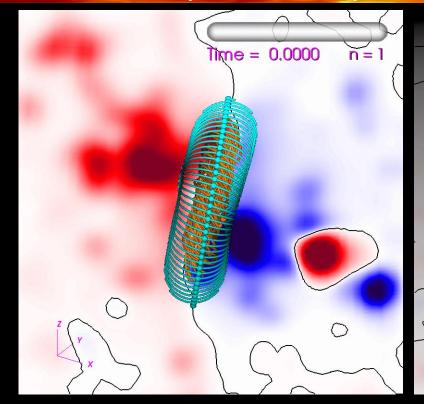
Acceptable time-to-solution: 20 min

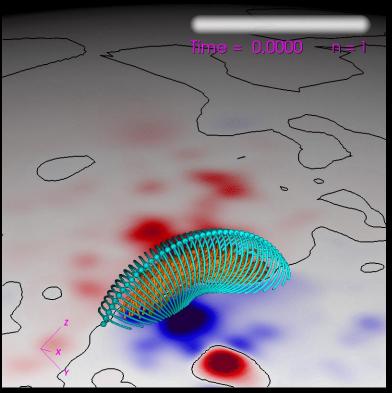
Detailed run information

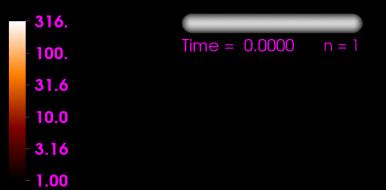
	N	Δ_{\min}	$\Delta_{ m max}$	max	$\frac{\Delta_{i+1} - \Delta_i}{\Delta_{i+1}}$
r	136	$64\mathrm{km}$	$540000\mathrm{km}$		9%
θ	450	0.052°	7.45°		11%
ϕ	543	0.052°	14.32°		10%
t	887	$0.001 \sec$	$0.13 \sec$		10%

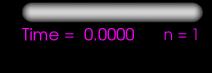
PCG Solver Iterations per Time Step (mean)

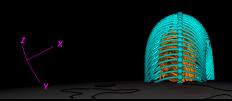
	SI Predictor	SI Corrector	Viscosity
PC1	186	195	963
PC2	$55 \rightarrow 65$	$57 \rightarrow 67$	$174 \rightarrow 310$











Production Test Run (TEST2)

TEST2: Eruptive Rope (high resolution)

Run information

Physical time duration: **118 sec**

Number of time-steps: 887

 $\overline{136 \times 450 \times 543} \sim 33 \, \mathrm{million \ points}$

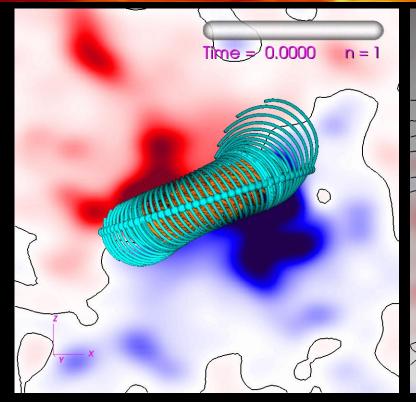
Acceptable time-to-solution: 90 min

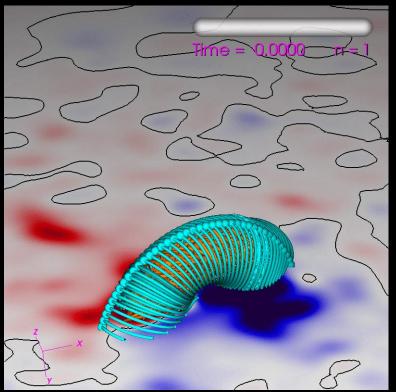
Detailed run information

	N	Δ_{\min}	$\Delta_{ m max}$	max	$\frac{\Delta_{i+1} - \Delta_i}{\Delta_{i+1}}$
r	160	$800\mathrm{km}$	$530000\mathrm{km}$		9%
θ	267	0.086°	9.74°		11%
$\overline{\phi}$	602	0.097°	14.90°		10%
t	200	$0.001 \sec$	$0.17 \sec$		4%

PCG Solver Iterations per Time Step (mean)

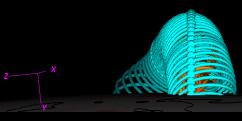
	SI Predictor	SI Corrector	Viscosity
PC1	32	34	516
PC2	$13 \to 15$	$14 \to 16$	$92 \rightarrow 149$











Motivation for OpenACC Implementation



- MAS run currently requires an HPC cluster for acceptable "time-to-solutions"
- Would rather run "in-house" to avoid wait queues, allocation usage, and have control of software stack







THE BIG IDEA: Can we achieve the same acceptable "time-to-solutions" on a single multi-GPU node using OpenACC in a *portable*, *single-source* implementation?



4xGPU Workstation



8xGPU Server



Server

DIFFUSE Recap (3.5 million pt test)

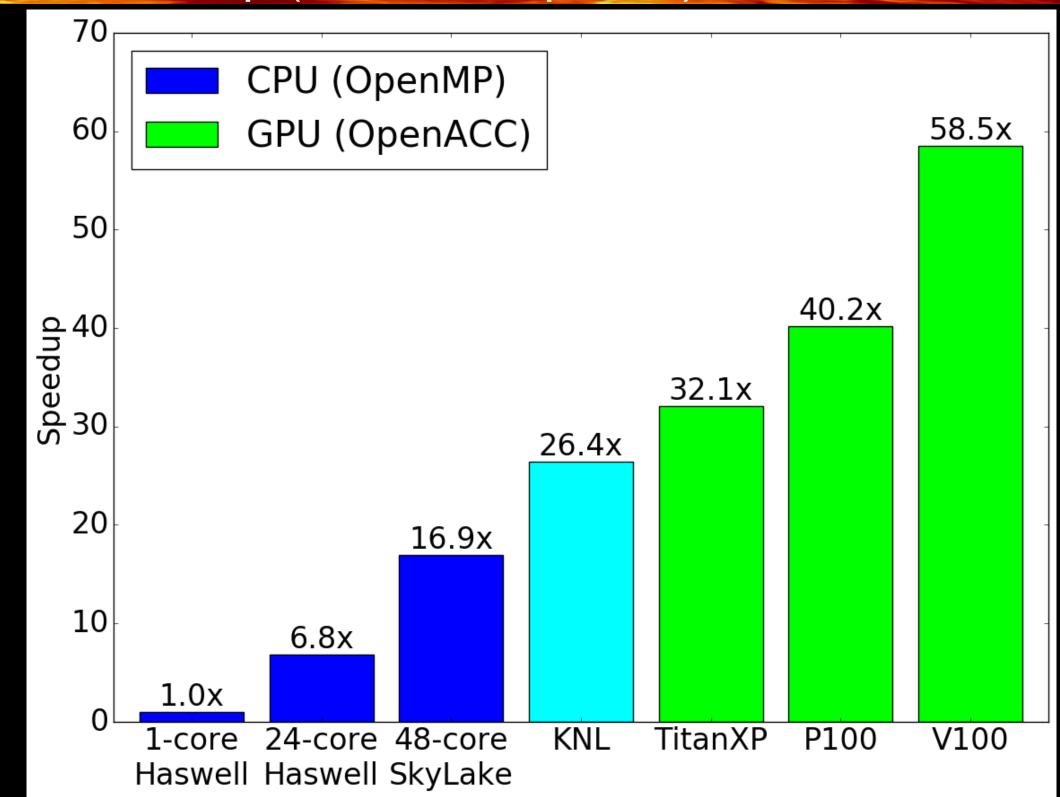


- Smooths unresolvable structure
- Integrates

$$\frac{\partial B_r}{\partial t} = \nabla_\perp^2 B_r$$

with explicit super time-stepping

Parallelized with OpenMP and OpenACC





POT3D Recap (200 million pt. test)



Solves potential field:

$$\nabla^2 \Phi = 0, \ \mathbf{B} = \nabla \Phi$$

- MPI+OpenACC
- PreconditionedConjugate Gradient

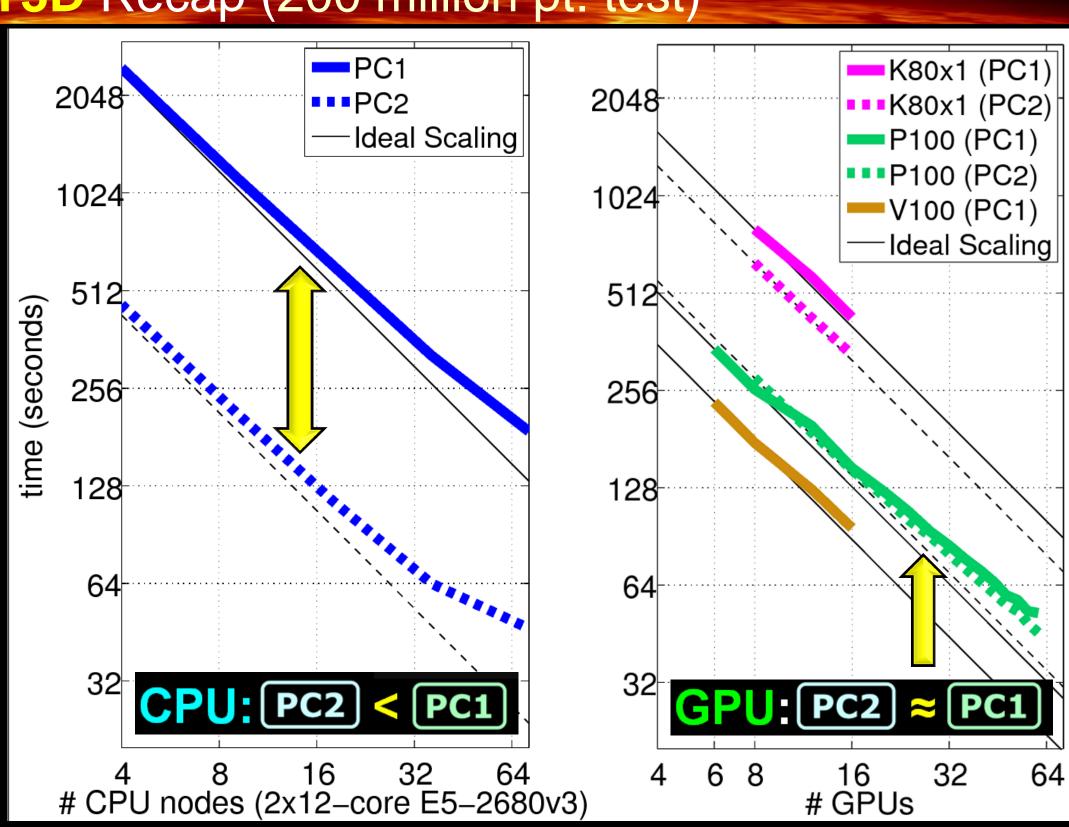
Two preconditioners:

Point-Jacobi
With ILU0

GPU Implementations:

PC1: pragmas only (portable)

PC2: cuSparse (not portable)

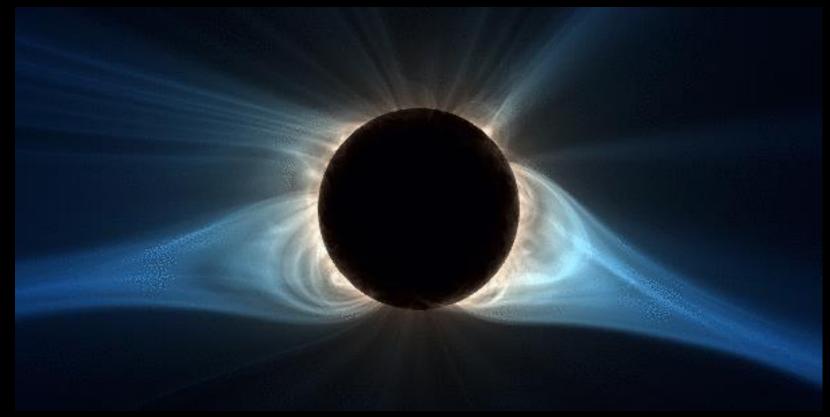




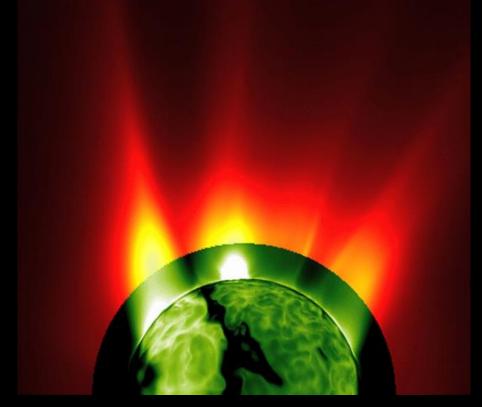


- Established MHD code with over 15 years of development used extensively in solar physics research
- Written in FORTRAN 90 (~50,000 lines), parallelized with MPI
- Available for use at the Community Coordinated Modeling Center (CCMC)





Predicted Corona of the August 21st, 2017 Total Solar Eclipse



Simulation of the Feb. 13th, 2009 CME

MAS: Full MHD Model Equations

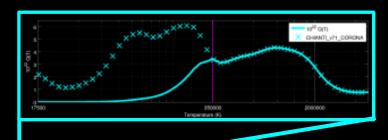
$$\frac{\partial \mathbf{A}}{\partial t} = \mathbf{v} \times (\nabla \times \mathbf{A}) - \left[\frac{c^2 \eta}{4 \pi} \nabla \times \nabla \times \mathbf{A} \right]$$

$$\frac{\partial \rho}{\partial t} = -\nabla \cdot (\rho \, \mathbf{v})$$

$$\frac{\partial T}{\partial t} = -\nabla \cdot (T\mathbf{v}) - (\gamma - 2)(T\nabla \cdot \mathbf{v}) + \frac{\gamma - 1}{2k} \frac{m_p}{\rho} \left[-\nabla \cdot (\mathbf{q}_1 + \mathbf{q}_2) \right]$$

$$\mathbf{q}_1 = -f(r)\,\beta_{\scriptscriptstyle{\mathrm{Tcut}}}(T)\,\kappa_0\,T^{5/2}\,\mathbf{\hat{b}}\mathbf{\hat{b}}\cdot\nabla T$$

$$\mathbf{q}_2 = (1-f(r))\,\frac{k}{(\gamma-1)}\,\frac{\rho}{m_p}\,T\,\mathbf{v}\,\mathbf{\hat{b}}\mathbf{\hat{b}}$$



COOLING

$$\left|\frac{\rho^2}{m_p^2}Q(T)\right| + H$$

CORONAL HEATING

$$H=H^*+rac{
ho}{4\,\lambda_\perp}\,\left[|z_-|\,z_+^2+|z_+|\,z_-^2
ight]
onumber \ \lambda_\perp=rac{\lambda_0}{|\mathbf{B}|}\,\sqrt{rac{B_w}{|\mathbf{B}|}}\,\,|z_\pm(r=R_\odot)|=oldsymbol{z}_0$$



$$\frac{\partial \epsilon_{\pm}}{\partial t} = -\nabla \cdot (\epsilon_{\pm} \left[\mathbf{v} \pm \mathbf{v}_{\mathbf{A}} \right]) - \frac{\epsilon_{\pm}}{2} \nabla \cdot \mathbf{v}$$

$$\frac{\partial \mathbf{v}}{\partial t} = -\mathbf{v} \cdot \nabla \mathbf{v} + \frac{1}{\rho} \left[\frac{1}{c} \mathbf{J} \times \mathbf{B} - \nabla p - \left[\nabla \left(\frac{\epsilon_{+} + \epsilon_{-}}{2} \right) \right] + \rho \mathbf{g} \right] + \left[\frac{1}{\rho} \nabla \cdot (\nu \rho \nabla \mathbf{v}) \right] + \left[\frac{1}{\rho} \nabla \cdot \left(S \rho \nabla \frac{\partial \mathbf{v}}{\partial t} \right) \right]$$







$$\frac{\partial z_{\pm}}{\partial t} = -\left(\mathbf{v} \pm \mathbf{v_A}\right) \cdot \nabla z_{\pm} - \frac{z_{\pm} |z_{\mp}|}{2 \lambda_{\perp}} + \frac{z_{\pm}}{4} \left(\mathbf{v} \mp \mathbf{v_A}\right) \cdot \nabla \left(\ln \rho\right) + \frac{z_{\mp}}{2} \left(\mathbf{v} \mp \mathbf{v_A}\right) \cdot \nabla \left(\ln |\mathbf{v_A}|\right)$$

MAS: MHD Model Equations ("Zero-Beta")

- In the low corona outside of active regions, the plasma beta is very small (i.e. dynamics dominated by magnetic field)
- Therefore, one can approximate the magnetic field and onset dynamics of the CME eruption with a simplified "zero-beta" form of the MHD equations

$$\frac{\partial \mathbf{A}}{\partial t} = \mathbf{v} \times (\nabla \times \mathbf{A}) - \frac{c^2 \eta}{4 \pi} \nabla \times \nabla \times \mathbf{A}$$

$$\frac{\partial \mathbf{v}}{\partial t} = -\mathbf{v} \cdot \nabla \mathbf{v} + \frac{1}{\rho} \left[\frac{1}{c} \mathbf{J} \times \mathbf{B} \right] + \frac{1}{\rho} \nabla \cdot (\nu \rho \nabla \mathbf{v}) + \frac{1}{\rho} \nabla \cdot \left(S \rho \nabla \frac{\partial \mathbf{v}}{\partial t} \right)$$
VISCOSITY

OF EXAMPLE 1

$$p = 0$$

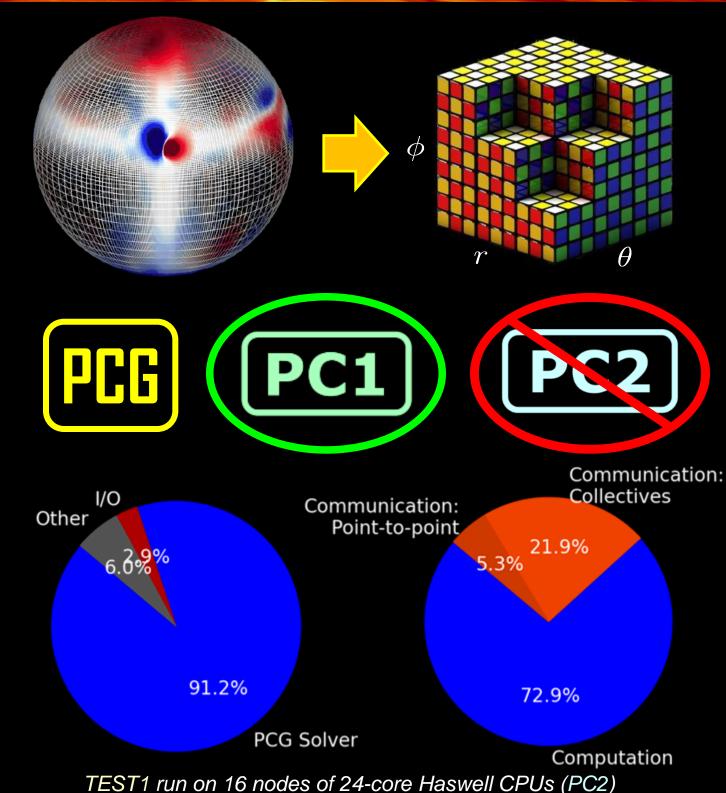
$$\rho = \rho_0(\mathbf{r})$$

$$\mathbf{B} = \nabla \times \mathbf{A} \qquad \nabla \cdot \mathbf{B} = 0 \qquad v_A^2 = |\mathbf{B}|^2 / (4\pi\rho) \qquad S = (\Delta t^2 \tilde{k}^2)^{-1} \left(C_w^2 / (1 - C_f)^2 - 1 \right) \qquad C_w^2 = 0.25 \, \Delta t^2 \, \tilde{k}^2 (v_c^2 + v_A^2)$$

$$\mathbf{J} = \frac{c}{4\pi} \nabla \times \mathbf{B} \quad \mathbf{g} = 0 \qquad C_f = \Delta t \, \tilde{k} \cdot \mathbf{v} \quad v_c = 0 \quad \tilde{k}^2 = 4 \left(\Delta r^{-2} + (r \, \Delta \theta)^{-2} + (r \, \Delta \phi \, \sin \theta)^{-2} \right)$$

MAS: Algorithm Summary and Profile

- Finite difference on non-uniform spherical grid
- Explicit and implicit time-step algorithms
- PCG used to solve implicit steps
- Sparse matrix operators stored in mDIA format, PC2 ILU0 matrix stored in CSR
- PCG solvers use the same PCs in POT3D. Since GPU results showed PC1~PC2, we only implement PC1 in MAS (portable!)
- PCG solvers are ~90% of run-time!



OpenACC Implementation: Quick Picks

Most implementation details the same as POT3D (see our GTC17 talk)





Managed Memory

Transition from managed memory to manual memory can be a BIG, all-or-nothing step

CPU Redundant Routines

Some calls use GPU, some don't.

OpenACC "if/if_present" conditional clauses to the rescue! (PGI > 18.1)

Valuable **PGI** ENVs

PGI_ACC_DEBUG
PGI_ACC_NOTIFY
PGI_ACC_TIME

PGI_ACC_PROFILE

PGI ACC FILL

GPU Data Residency

Avoiding GPU-CPU data transfers can involve increased development time due to many small (possibly awkward) routines

OpenACC Implementation: Derived Types

Fortran Derived Types

```
type :: vvec
    real, dimension(:,:,:), allocatable :: r
    real, dimension(:,:,:), allocatable :: t
    real, dimension(:,:,:), allocatable :: p
end type
type :: vvec_bc
    type(vvec) :: r0
    type(vvec) :: r1
end type
```

```
type(vvec),target :: v
type(vvec_bc),target :: v_bc
(Allocations...)
```

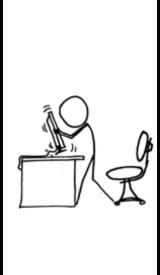
```
!$acc parallel loop collapse(2)
!$acc& default(present)
do j=2,ntm1
    do i=2,nrm-1
        v%r(i,j,2)=v%r(i,j,2)+v_bc%r0%r(i,j,2)
    enddo
enddo
```

```
"Manual" Deep-copy
  !$acc enter data create(v,v%r,v%t,v%p)

  !$acc enter data create(
  !$acc& v_bc,v_bc%r0,v_bc%r1,
  !$acc& v_bc%r0%r,v_bc%r0%t,v_bc%r0%p,
  !$acc& v_bc%r1%r,v_bc%r1%t,v_bc%r1%p)

"True" Deep-copy (PGI: -ta:tesla,deepcopy)
  !$acc enter data create(v)
  !$acc enter data create(v_bc)
```

- "True" Deep-copy + CUDA-aware MPI weren't playing nicely, so we used manual deep-copy
- Due to compiler bug (fixed in PGI
 ≥17.10), had to change code to only
 use single-level types
- Due to compiler bug (PGI ≥17.10) with CUDA-aware MPI + types, used PGI 17.9 (work-around found)



OpenACC Implementation: Array Reductions

Array Reductions

OpenACC scalar reductions

```
real(r_typ) :: sum
!$acc kernels loop
!$acc& reduction(+:sum)
do j=1,m
   sum=sum+a(j)
enddo
```

OpenACC does *not* directly support array reductions

```
allocate(sum(n))
do j=1,m
  sum(:)=sum(:)+a(:,j)
enddo
```

Two example options

(1)

```
!$acc kernels
!$acc loop
do j=1,m
!$acc loop
    do i=1,n
!$acc atomic update
        sum(i)=sum(i)+a(i,j)
enddo
```

Timing results of 1 step of TEST1 on TitanXP

	GPU	CPU	CPU
			(full routine)
(1)	0.25	2.3	21.3
$\overline{(2)}$	0.46	2.5	17.6

(2)

```
!$acc kernels loop
do i=1,n
   sum(i)=SUM(a(i,1:m))
enddo
```

We use option (2) for code simplicity

Full routine only 0.03% of total run time

For TEST1, ~60% of wall-time in computing velocity matrix multiply routine

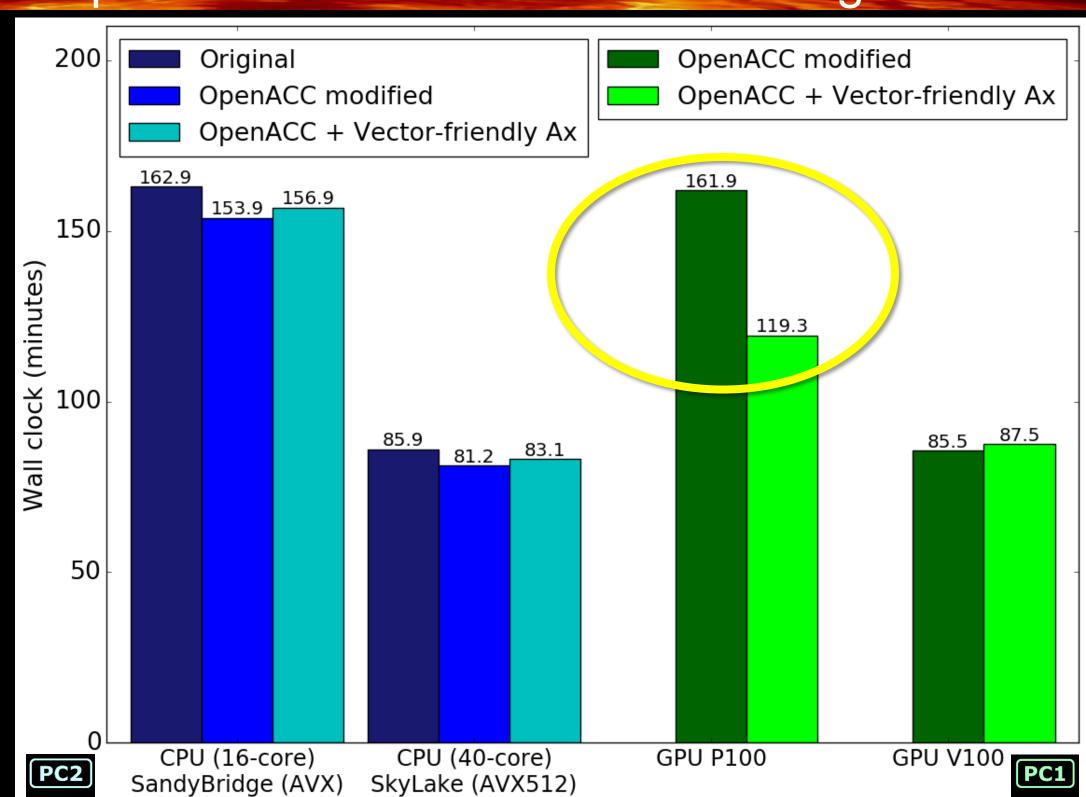
Cache-friendly vs Vector-friendly

```
do k=2, npm1
do j=2, ntm1
  do i=2,nrm-1
   ii=ntm2*(nrm-2)*(k-2)
      +(nrm-2)*(j-2)+(i-1)
   q(ii)=a_r(1,1,j,k)*ps%r(i ,j ,k-1)
        +a_r(2,i,j,k)*ps%r(i,j-1,k)
        +a_r<mark>(14,</mark>i,j,k)*ps%p(i ,j ,k )
        +a_r(15,i,j,k)*ps%p(i+1,j,k)
  enddo
 enddo
enddo
```

```
do k=2, npm1
 do j=2, ntm1
  do i=2, nrm-1
   ii=ntm2*(nrm-2)*(k-2)
      +(nrm-2)*(j-2)+(i-1)
  q(ii)=a_r( i,j,k<mark>,1)</mark>*ps%r(i ,j ,k-1)
        +a_r(i,j,k,2)*ps%r(i,j-1,k)
        +a_r(i,j,k,14)*ps%p(i,j,k)
        +a_r(i,j,k<mark>,15)</mark>*ps%p(i+1,j ,k )
  enddo
 enddo
enddo
```

Cache vs Vector Results (TEST1)

- CPU: Vectorfriendly version slower, but still faster than original code
- GPU: Vectorfriendly version much faster on P100, little change on V100





```
parallel, kernels, gangs, workers, vectors ... oh my!
```

Many configuration options (hardware narrows choices a bit)

```
(kernels)
                                                          vr,vt,vp computed
!$acc parallel default(present) present(ps) async(1)
                                                          asynchronously
!$acc loop
     do k=2, npm1
                            num_gangs(#), num_workers(#), num_vectors(#)
!$acc loop
                            gang, worker(#), vector(#), seq
       do j=2, ntm1
!$acc loop
          do i=2,nrm-1
```

We test various clause options with 1 step of TEST1 on a TitanXP GPU (timing routine using PGI_ACC_TIME=1)

Source	PGI 17.9 Output	Time (s)
<pre>!\$acc parallel !\$acc loop !\$acc loop !\$acc loop</pre>	<pre>!\$acc loop gang ! blockidx%x !\$acc loop seq !\$acc loop vector(128) ! threadidx%x</pre>	60.3
<pre>!\$acc parallel vector_length(32) !\$acc loop !\$acc loop !\$acc loop</pre>	<pre>!\$acc loop gang ! blockidx%x !\$acc loop seq !\$acc loop vector(32) ! threadidx%x</pre>	55.7
<pre>!\$acc parallel vector_length(16) !\$acc loop !\$acc loop !\$acc loop</pre>	<pre>!\$acc loop gang ! blockidx%x !\$acc loop seq !\$acc loop vector(16) ! threadidx%x</pre>	76.8
<pre>!\$acc loop independent !\$acc loop independent !\$acc loop independent</pre>	<pre>!\$acc loop gang ! blockidx%y !\$acc loop gang, vector(4) ! blockidx%z threadidx%y !\$acc loop gang, vector(32) ! blockidx%x threadidx%x</pre>	45.7
<pre>!\$acc loop independent gang worker vector !\$acc loop independent gang worker vector !\$acc loop independent gang worker vector</pre>	<pre>!\$acc loop gang ! blockidx%z !\$acc loop gang, vector(4) ! blockidx%y threadidx%z !\$acc loop gang, worker(2), vector(64) ! blockidx%x threadidx%y threadidx%x</pre>	47.7
<pre>!\$acc loop independent gang !\$acc loop independent gang worker !\$acc loop independent gang vector</pre>	<pre>!\$acc loop gang ! blockidx%z !\$acc loop gang, worker(4) ! blockidx%y threadidx%y !\$acc loop gang, vector(32) ! blockidx%x threadidx%x</pre>	49.1
<pre>!\$acc loop independent gang !\$acc loop independent vector(8) !\$acc loop independent vector(8)</pre>	<pre>!\$acc loop gang ! blockidx%x !\$acc loop gang, vector(8) ! blockidx%z threadidx%y !\$acc loop gang, vector(8) ! blockidx%y threadidx%x</pre>	151.84
<pre>!\$acc loop independent gang !\$acc loop independent gang vector(8) !\$acc loop independent gang vector(8)</pre>	<pre>!\$acc loop gang ! blockidx%x !\$acc loop gang, vector(8) ! blockidx%z threadidx%y !\$acc loop gang, vector(8) ! blockidx%y threadidx%x</pre>	91.93

OpenACC Implementation: Effort Summary

 $\approx 1\%$

OpenACC comment lines

≈8%

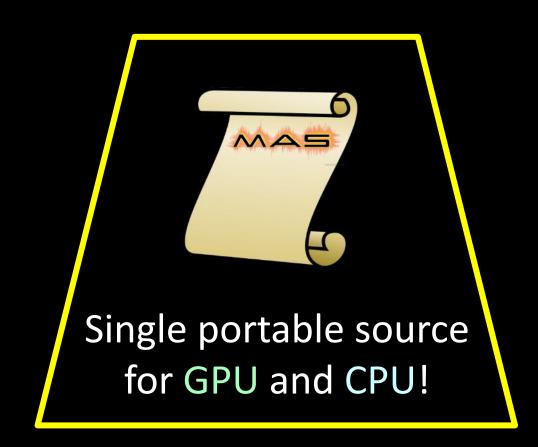
Total added, deleted, and changed lines

Details

Total lines in original code	51,591
Total lines in accelerated code	54,191
Total !\$acc/!\$acc& lines added	671 (1.0%)
Total modified lines	844 (1.6%)
Total # of additional lines	2600 (5.0%)
Total # of different lines	4314 (8.0%)

Factors to consider

- Added Lines: Duplicate CPU routines (can remove with OpenACC 2.6 conditionals)
- Deleted Lines: Optional CPU code simplifications
- Modified Lines: CPU changes for array reductions, vector-friendly matrix multiply, and single-level derived types (temporary)
- OpenACC Comment Lines: Full code not accelerated (zero-beta parts only!)



Timing Procedures

- "Time-to-solution" includes
 I/O, comm, setup, etc.
 (Queue times excluded, but important!)
- Acceptable "time-to-solution" for TEST1 & TEST2 set by current pipeline (not cherry picked!)
- We use best available compiler, compiler version, instruction sets, library versions, and algorithm for each system

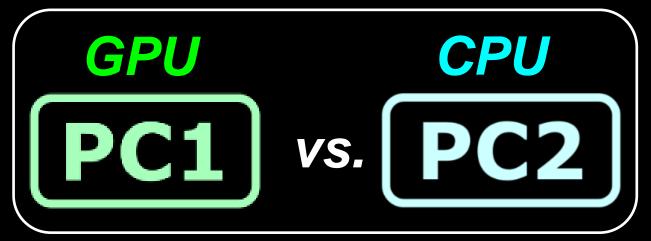
Why is this fair?

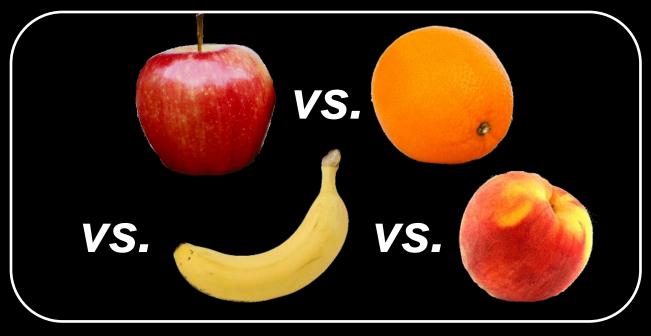
We care about the "real" world

We are not benchmarking hardware

We want to test the maximum performance on each system for solving our problem, using our code







Hardware and Environments

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	NASA NAS Pleiades & Electra					SDSC Comet	TACC St	tampede2
Compiler	Intel 2018 .0.128					Intel 2016.3.210	Inte	el 18.0.0
MPI Library	SGI MPT 2.15r20					MVAPICH2 v2.1	Intel N	ИРІ 18.0.0
Intel Family	Sandy Bridge	Ivy Bridge	Haswell	Broadwell	Skylake	Haswell	KNL	Skylake
Instruction Set	AVX	AVX	AVX2	AVX2	AVX512	AVX2	AVX512	AVX512
Processor	E5-2670	E5-2680v2	E5-2680v3	E5-2680v4	Gold 6148	E5-2680v3	Phi 7250	Platinum 8160
Clock Rate	2.6 GHz	2.8 GHz	2.5 GHz	2.4 GHz	2.4 GHz	2.5 GHz	1.4 GHz	2.1 GHz
# Cores	16	20	24	28	40	24	68	48
Memory Bandwidth	51.2 GB/s	51.2 GB/s 59.7 GB/s 68 GB/s 76.8 GB/s 128 GB/s		68 GB/s	115.2 GB/s	128 GB/s		



	NVIDI	SDSC Comet	
Compiler	PGI	17.9	PGI 17.10
MPI Library	OpenMi	인 1.10.7	OpenMPI 2.1.2
CUDA Library	CUDA 9	9.0.176	CUDA 8.0
Driver Version	387	367.48	
Model (# GPUs/node)	P100 PCIE (4)	V100 PCIE (4)	P100 PCIE (4)
Compute Capability	6.0 7.0		6.0
Clock Rate	1.33 GHz 1.38 GHz		1.33 GHz
# CUDA DP Cores	1792 2560		1792
Memory Bandwidth	732 GB/s	900 GB/s	732 GB/s









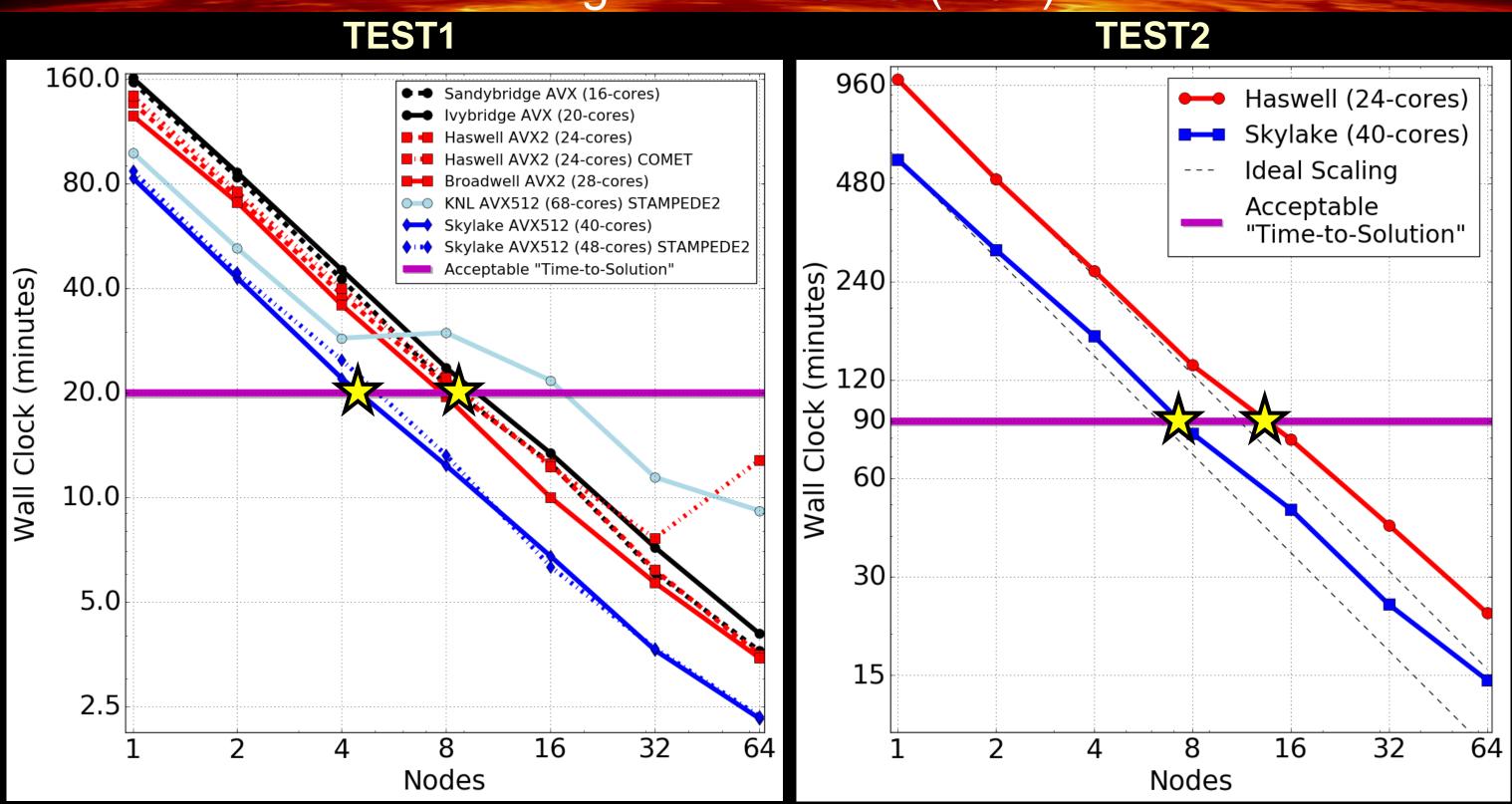




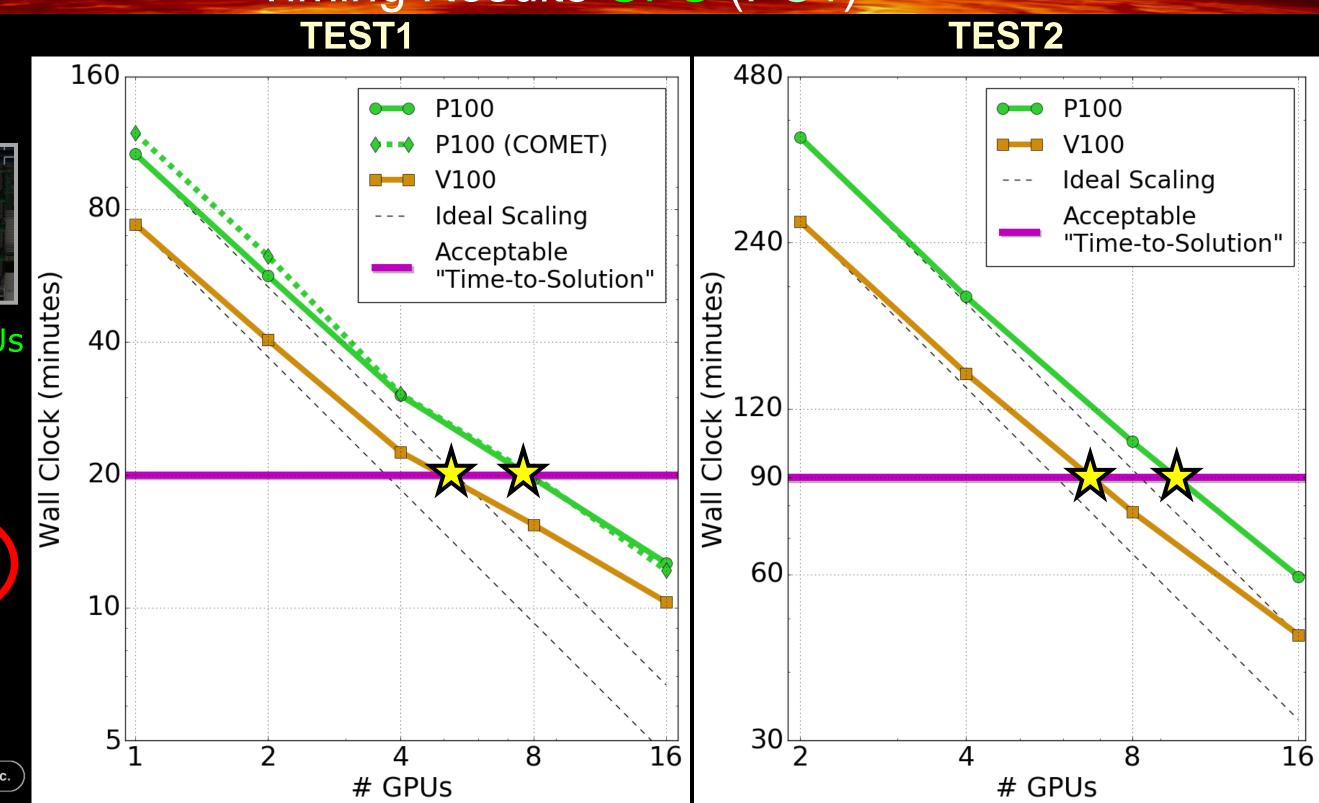
Extreme Science and Engineering Discovery Environment



Timing Results CPU (PC2)



Timing Results GPU (PC1)





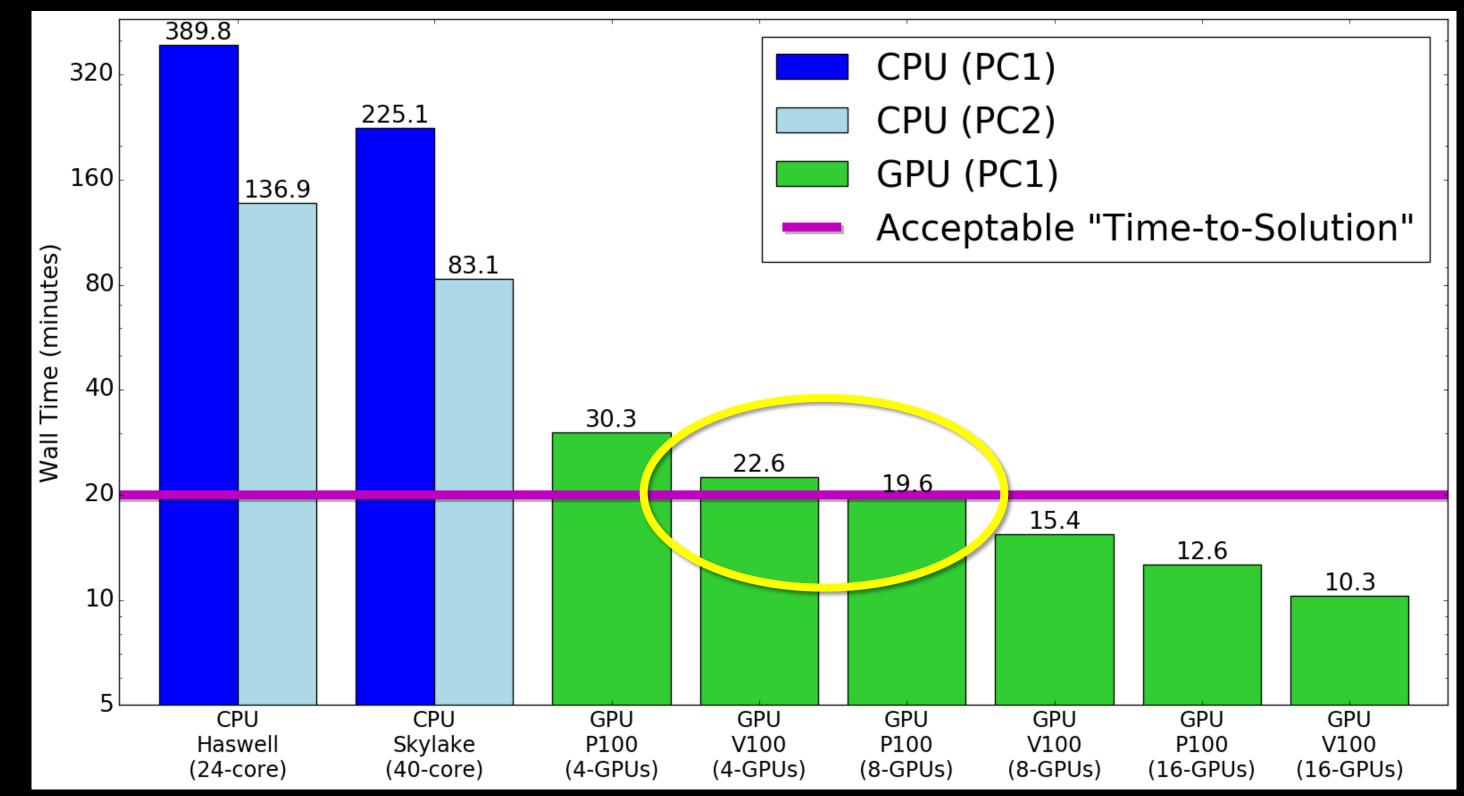
4x PCIe GPUs per node



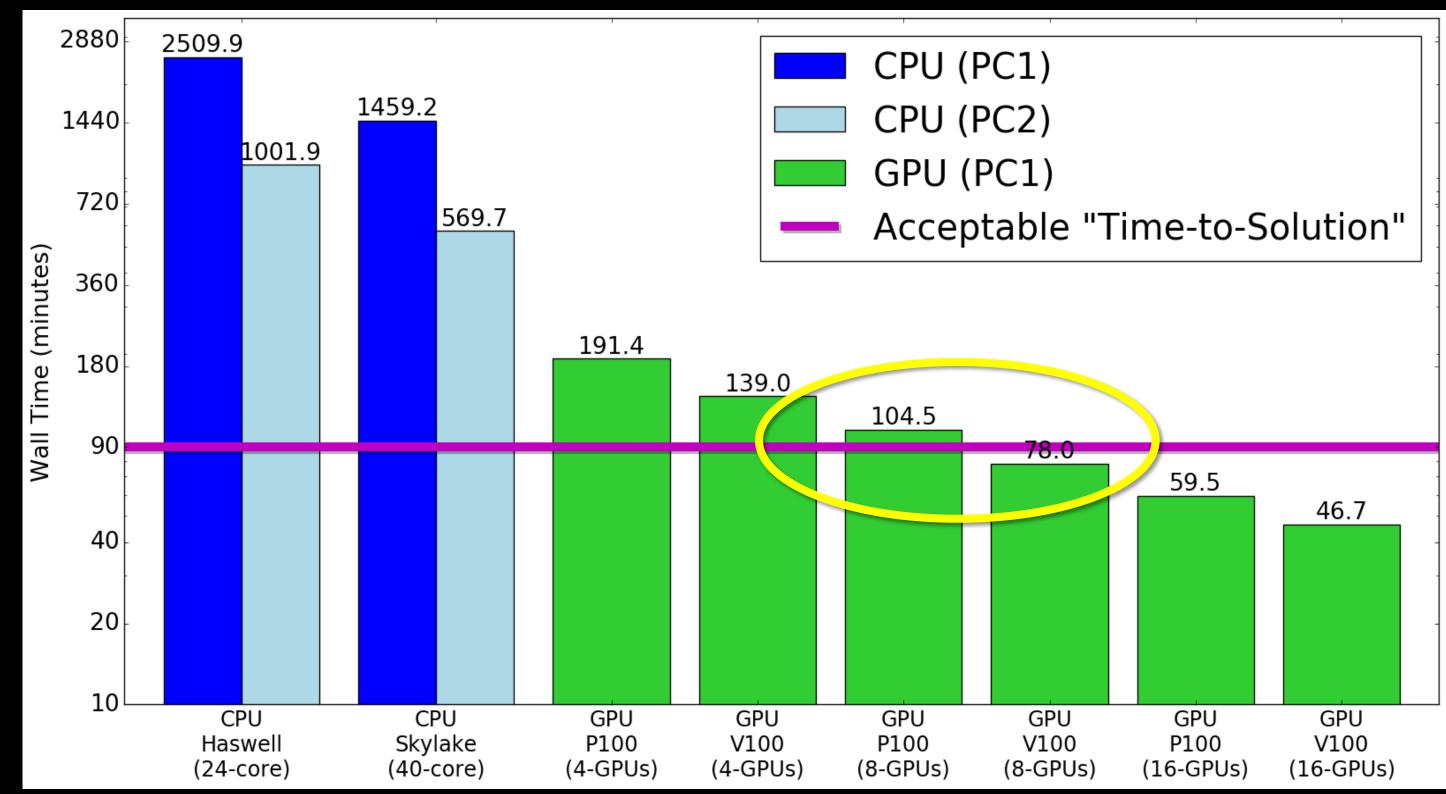


Predictive Science Inc.

Timing Results Single Node ("In-house")



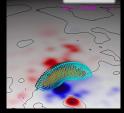
Timing Results Single Node ("In-house")



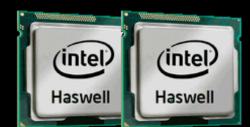
Performance Summary

TEST1:

Acceptable time-to-solution: 20 min







2x12-core Haswell Nodes



 $\approx 4x$



2x20-core Skylake Nodes



















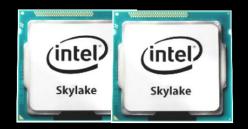








2x12-core Haswell Nodes



2x20-core Skylake Nodes







Summary and Outlook

THE BIG IDEA:

Can we achieve the same acceptable "time-tosolutions" on a single multi-GPU node using OpenACC in a portable, single-source implementation?





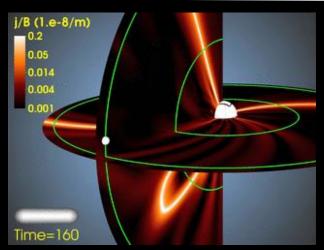


Yup!

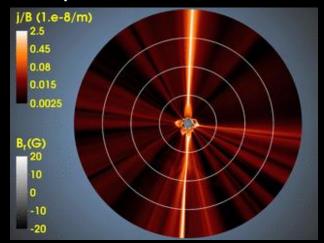
4xGPU Workstation

For TEST1 and TEST2 (representative of many cases), we can move from HPC cluster to "in-house"!

- Future improvements (Give PC2 another go? Mixed-precision?)
- Next steps in OpenACC implementation of MAS:
 - Heliospheric runs (where PC1 is most efficient on the CPU runs)
 - Thermodynamic runs (Using many multiple-GPU nodes)



Heliospheric CME Simulation



Thermodynamic CME Simulation

Questions?



OpenACC User Group

Twitter @OpenACCorg Facebook @OpenACCorg LinkedIn OpenACC Developers



- NSF's Frontiers in Earth System Dynamics program
- NASA's Living with a Star program
- Air Force Office of Scientific Research

We gratefully acknowledge NVIDIA Cooperation for donating allocation use of their PSG Cluster for GPU timings.











